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Leading nucleon and the proton-nucleus inelasticity

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Abstract

We present in this paper, a calculation of average proton-nucleus inelasticity. Using an iteractive leading particle model and the Glauber model, we relate the leading particle distribution in nucleon-nucleus iteractions with the respective one in nucleon-proton collisions. To describe the leading particle distribution in nucleon-proton, we use the Regge-Mueller formalism.

We calculate the average proton-nucleus inelasticity. Using an Iterative Leading Particle Model (Frichter et al., 1997) and the Glauber model (Glauber, 1959; Glauber et al., 1970), we relate the leading particle distribution in nucleon-nucleus interactions with the respective one in nucleon-proton collisions. In this model the leading particle spectrum in $p+A \to N(\text{nucleon})+X$ collisions is built from successive interacions with ν interacting proton of the nucleus A and the behaviour is controlled by a straightforward convolution equation. It should be mentioned that, strictly speaking, the convolution should be 3-dimensional. Here we only considered the 1-dimension approximation. In a recent paper (Bellandi et al., 1999) we have used this model to describe the hadronic flux in the atmosphere, showing that the average nucleon-nucleus elasticity, $\langle x \rangle_{N-p}$, by means of the following relation

$$(1 - \langle x \rangle_{N-A}) = \frac{1}{\sigma_{in}^{N-ar}} \int d^2b \left[1 - \exp[-(1 - \langle x \rangle_{N-p})\sigma_{tot}^{pp} T(b)] \right]$$
 (1)

where T(b) is the nuclear thickness and given by means of the Woods-Saxon model (Woods, & Saxon, 1954; Barrett, & Jackson, 1977). Introducting the inelasticity given by < k >= 1 - < x >, this expression can be transformed in

$$< k>_{N-A} = \frac{1}{\sigma_{in}^{N-A}} \int d^2b \left[1 - \exp[- < k>_{N-p} \sigma_{tot}^{pp} T(b)] \right].$$
 (2)

It is clear from this relationship that only in small σ^{pp}_{tot} limit is $< K>_{N-A} \simeq < K>_N$. In general, $< K>_{N-A} \geq < K>_N$ and the effect increasing with the increase of σ^{pp}_{tot} . If $< K>_{N} \rightarrow 0$, one also has $< K>_{N-A} \rightarrow 0$. On the other hand, if $< K>_{N} = 1$, then $< K>_{N-A} = 1$, and Eq. (2) reproduces the Glauber model relationship between σ^{N-A}_{in} and σ^{pp}_{tot} . In order to calculate $< K>_{N-A}$ we use for $< K>_N$ the values calculated by means of the Regge-Mueller formalism (Batista et al., 1998) and as input for σ^{pp}_{tot} we have used the UA4/2 parametrizations for the energy dependence (Burnett et al., 1992). In the Fig. (1) we show the results of this calculations for the following nuclei: C, Al, Cu, Ag, Pb and air (A=14.5). In this figure we also show recent experimental data for p-Pb, $< K> = 0.84 \pm 016$ (Barroso et al., 1997) and for p-C, $< K> = 0.65 \pm 0.08$ (Wilk & Wlodarczyk, 1999).

In the Fig. (2), we compare the calculated $< K >_{p-air}$ with results from some models used in Monte Carlo simulation (Gaisser et al., 1993); the Kopeliovich *et al.* (Kopeliovich et al., 1989) (KNP) QCD multiple Pomeron exchanges model; the Dual Parton model with sea-quark interaction of Capella *et al.* (Capella et al., 1981); the statistical model of Fowler *et al.* (Fowler et al., 1987) and with calculated values derived from cosmic ray data by Bellandi *et al.* (Bellandi et al., 1998). We note that the calculated $< K >_{p-air}$ (Bellandi et al., 1998) was done assuming for the T(b) nuclear thickness the Durand and Pi model (Durand & Pi, 1988), which gives small values for the average inelasticity. In the Fig. (2) we also show the average inelasticity values as calculated by means of this model.

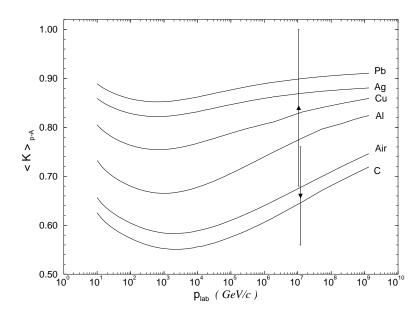


Figure 1: Proton-nucleus inelasticities calculated. The up triangle is Pb data (Barroso et al., 1997) and down triangle is C data (Wilk & Wlodarczyk, 1999).

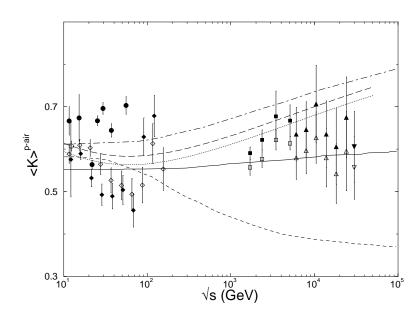


Figure 2: The $< K >^{p-air}$ as a function of \sqrt{s} in GeV. The experimental data from (Bellandi et al., 1998). Dash line from (Fowler et al., 1985). Solid line from (Capella et al., 1981). Dot-dash line from (Kopeliovich et al., 1989). Dot line from Eq. (2) with Woods-Saxon model (Woods & Saxon, 1954; Barrett, & Jackson, 1977). Long-dash line from Durand-Pi model (Durand & Pi, 1988).

References

Barrett, R.C. & Jackson, D.F. 1977, Nuclear Sizes and Structure, (Clarendon Press, Oxford).

Barroso, S.L.C. et al. (Chacaltaya Collab.) 1997, Proc. 25th ICRC, (Durban, 1997)

Batista, M. & Covolan, R.J.M. 1998, hep-ph:9811425.

Bellandi, J., Fleitas, J.R. & Dias de Deus, J. 1998, Il Nuovo Cimento, 111A, 149.

Bellandi, J. et al. 1999, Proc. 26th ICRC, (Salt Lake City, 1999).

Burnett, T.H. et al., (UA4/2 Collab.) 1992, ApJ 349 L25.

Capella, A. et al. 1981, Z. Phys. C 10, 249.

Durand, L. & and Pi, H. 1988, Phys. Rev. **D** 38, 78.

Fowler, G. et al. 1987, Phys. Rev. D 35, 870.

Frichter, G.M. et al. 1997, Phys. Rev. **D** 56, 3135.

Gaisser, T.K. et al. 1993, Phys. Rev. D 47, 1919.

Glauber, R.J. 1959, Lect. Theor. Phys. Vol.1, edited by W.Britten and L.G.Dunhan (Interscience, NY), 135.

Glauber, R.J. et al. 1970, Nucl. Phys.B 12, 135.

Kopeliovich, B.Z. et al. 1989, Phys. Rev. **D** 39, 769.

Wilk, G. & Wlodarczyk, Z. 1999, Phys. Rev. **D** 57, 180.

Woods, R.D. & Saxon, D.S. 1954, Phys. Rev. 95, 577.

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